

### 234(3): Minkowski Metric; Relativistic Dynamics.

The relativistic dynamics mean that time derivatives are derivatives with respect to proper time.

Therefore:

$$\underline{r} = r \underline{e}_r \quad - (1)$$

$$\underline{v} = \left( \frac{dr}{d\tau} \right) \underline{e}_r + r \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta \quad - (2)$$

The unit vectors obey the equations:

$$\frac{d\underline{e}_r}{d\tau} = \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta \quad - (3)$$

$$\frac{d\underline{e}_\theta}{d\tau} = - \left( \frac{d\theta}{d\tau} \right) \underline{e}_r \quad - (4)$$

The acceleration is therefore:

$$\underline{a} = \frac{d\underline{v}}{d\tau} = \frac{d}{d\tau} \left( \left( \frac{dr}{d\tau} \right) \underline{e}_r + r \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta \right)$$

$$\begin{aligned} &= \frac{d}{d\tau} \left( \frac{dr}{d\tau} \right) \underline{e}_r + \left( \frac{dr}{d\tau} \right) \left( \frac{d\underline{e}_r}{d\tau} \right) \\ &\quad + \left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta + r \frac{d}{d\tau} \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta + r \left( \frac{d\theta}{d\tau} \right) \left( \frac{d\underline{e}_\theta}{d\tau} \right) \end{aligned} \quad - (5)$$

Now use eqs. (3) and (4) in eq. (5):

$$\begin{aligned}
 \underline{a} &= \left( \frac{d^2 r}{d\tau^2} \right) \underline{e}_r + \left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta \\
 &+ \left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta + r \frac{d}{d\tau} \left( \frac{d\theta}{d\tau} \right) \underline{e}_\theta - r \left( \frac{d\theta}{d\tau} \right)^2 \underline{e}_r \\
 &= \left( \frac{d^2 r}{d\tau^2} - r \left( \frac{d\theta}{d\tau} \right)^2 \right) \underline{e}_r \quad - (6) \\
 &\quad + \left( r \frac{d^2 \theta}{d\tau^2} + 2 \left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) \right) \underline{e}_\theta
 \end{aligned}$$

From note 234(2):

$$\frac{d^2 r}{d\tau^2} \neq 0 \quad - (7)$$

but

$$\underline{p} = m \frac{dr}{d\tau} \quad - (8)$$

is a constant of motion. So: - (9)

$$\underline{a} = \left[ -r \left( \frac{d\theta}{d\tau} \right)^2 \right] \underline{e}_r + \left[ r \frac{d^2 \theta}{d\tau^2} + 2 \left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) \right] \underline{e}_\theta$$

The Mikushki force is:

$$\underline{F} = m \underline{a} \quad - (10)$$

3)

Here:

$$L = m r^2 \frac{d\theta}{d\tau} \quad - (11)$$

and

$$p = m \frac{dr}{d\tau} \quad - (12)$$

wt

$$E = m c^2 \frac{dt}{d\tau} \quad - (13)$$

In the Michuski theory of cosmology,  $E$ ,  $p$  and  $L$  are all constants of motion.

So: 
$$\left( \frac{d\theta}{d\tau} \right)^2 = \left( \frac{L}{m r^2} \right)^2 \quad - (14)$$

$$\frac{d}{d\tau} \left( \frac{d\theta}{d\tau} \right) = \frac{L}{m} \frac{d}{d\tau} \left( \frac{1}{r^2} \right) \quad - (15)$$

$$\left( \frac{dr}{d\tau} \right) \left( \frac{d\theta}{d\tau} \right) = \frac{p}{m} \cdot \frac{L}{m r^2} = \left( \frac{pL}{m^2} \right) \frac{1}{r^2} \quad - (16)$$

So:

$$\frac{a}{r} = - \left( \frac{L}{m} \right)^2 \frac{1}{r^3} \underline{e}_r + \frac{1}{m} \frac{dp}{d\tau} \underline{e}_r \quad - (17)$$

$$+ \left( \frac{L r}{m} \frac{d}{d\tau} \left( \frac{1}{r^2} \right) + \frac{2 p \cdot L}{m^2} \frac{1}{r^2} \right) \underline{e}_\theta$$

Finally use:

$$4) \quad \frac{df}{d\tau} = \frac{df}{dr} \frac{dr}{d\tau}, \quad f = \frac{-1}{r^2} \quad - (18)$$

$$\text{So} \quad \frac{d}{d\tau} \left( \frac{1}{r^2} \right) = -\frac{2}{r^3} \frac{dr}{d\tau} = -\frac{2p}{m r^3} \quad - (19)$$

$$\text{So:} \quad \underline{a} = - \left( \frac{L}{m} \right)^2 \frac{1}{r^3} \underline{e}_r + \frac{1}{m} \frac{dp}{d\tau} \underline{e}_r \quad - (20)$$

$$+ \left( \frac{2pL}{m^2} \frac{1}{r^2} - \frac{2pL}{m^2 r^2} \right) \underline{e}_\theta$$

i.e.

$$\underline{a} = \left[ \frac{1}{m} \frac{dp}{d\tau} - \left( \frac{L}{m} \right)^2 \frac{1}{r^3} \right] \underline{e}_r \quad - (21)$$

The Michowski force is:

$$\underline{F} = \left( \frac{1}{m} \frac{dp}{d\tau} - \frac{L^2}{m r^3} \right) \underline{e}_r \quad - (22)$$

Conclusion

The force (22) is equivalent to the infinitesimal line element:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dr^2 - r^2 d\theta^2 \quad - (23)$$

Eq. (22) is the acceleration (21) multiplied